Spectroscopy of ⁹⁹Cd and ¹⁰¹In: challenges and new approaches

Joochun (Jason) Park CENS Lunch Seminar Aug. 26, 2020





Preface

PHYSICAL REVIEW C 102, 014304 (2020)

Spectroscopy of ⁹⁹Cd and ¹⁰¹In from β decays of ⁹⁹In and ¹⁰¹Sn

J. Park , ^{1,2,*} R. Krücken, ^{1,2} A. Blazhev, ³ D. Lubos, ^{4,5,6} R. Gernhäuser, ⁴ M. Lewitowicz, ⁷ S. Nishimura, ⁵ D. S. Ahn, ⁵ H. Baba, ⁵ B. Blank, ⁸ P. Boutachkov, ⁹ F. Browne, ^{5,10} I. Čeliković, ^{7,11} G. de France, ⁷ P. Doornenbal, ⁵ T. Faestermann, ^{4,6} Y. Fang, ¹² N. Fukuda, ⁵ J. Giovinazzo, ⁸ N. Goel, ⁹ M. Górska, ⁹ H. Grawe, ⁹ S. Ilieva, ¹³ N. Inabe, ⁵ T. Isobe, ⁵ A. Jungclaus, ¹⁴ D. Kameda, ⁵ G. D. Kim, ¹⁵ Y.-K. Kim, ^{15,16} I. Kojouharov, ⁹ T. Kubo, ⁵ N. Kurz, ⁹ Y. K. Kwon, ¹⁵ G. Lorusso, ⁵ K. Moschner, ³ D. Murai, ⁵ I. Nishizuka, ¹⁷ Z. Patel, ^{5,18} M. M. Rajabali, ¹ S. Rice, ^{5,18} H. Sakurai, ^{5,19} H. Schaffner, ⁹ Y. Shimizu, ⁵ L. Sinclair, ^{5,20} P.-A. Söderström, ⁵ K. Steiger, ⁴ T. Sumikama, ¹⁷ H. Suzuki, ⁵ H. Takeda, ⁵ Z. Wang, ¹ H. Watanabe, ²¹ J. Wu, ^{5,22} and Z. Y. Xu¹⁹

DOI: 10.1103/PhysRevC.102.014304

Acknowledgements

••••

(Grant No. 25247045) of Japan Society for the Promotion of Science (JSPS). <u>This work was supported by the Institute</u> for Basic Science (IBS-R031-D1). The authors acknowledge also the support of the DFG cluster of excellence "Origin and Structure of the Universe," German BMBF under Contracts No. 05P15PKFNA and No. 05P19PKFNA, and



Most figures in this talk are taken from this article unless otherwise specified

Introduction



First level schemes of ⁹⁹Cd and ¹⁰¹In



M. Lipoglavšek et al., PRC 66, 011302(R) (2002)

Fusion evaporation vs decay spectroscopy



grazing reactions \rightarrow high-spin states populated

 $\Delta J \leq 1$, $\Delta \pi = 0$ states

Population of different states in residual/daughter nuclei

 $\beta\gamma$ spectra (⁹⁹In \rightarrow ⁹⁹Cd), this work



γγ coincidences and level scheme of ⁹⁹Cd



Comparing experiment and theory

In β -decay and γ -ray spectroscopy experiments for nuclear structure, the ideal observables are transition strengths. For Gamow-Teller decays:



The main difficulty lies with measurements of t_i , the partial decay half-lives (see next)

8

The Pandemonium effect

Volume 71B, number 2

PHYSICS LETTERS

21 November 1977

THE ESSENTIAL DECAY OF PANDEMONIUM: A DEMONSTRATION OF ERRORS IN COMPLEX BETA-DECAY SCHEMES

J.C. HARDY *, L.C. CARRAZ, B. JONSON ‡ and P.G. HANSEN ‡ CERN, Geneva, Switzerland

Monte-Carlo simulation of decay of a hypothetical nucleus "Pandemonium" with controllable inputs revealed flaws in $\beta\gamma$ experiments

Finite γ -ray detection efficiency and small branching ratios inevitable \rightarrow Biased beta-decay intensities of the strongest branches, and thus B_{GT}

Possible remedy via total absorption spectroscopy (TAS) with higher sensitivity and efficiency

Challenges of TAS: background contamination, energy resolution, etc

Extrapolating theory down to $I_{\beta v}$

Part A:

Theoretical B_{GT} values + mass difference (Q_{FC}) + final states' E_x \rightarrow T_{1/2} value, b_b distribution

Some of the masses of the most exotic isotopes have to be taken from SM, mass models, etc

Uncertainty estimates using RMS deviation of final state energies of models, typically on the order of 100 keV or 10% for empirical SM on the most exotic nuclei

Part B:

Theoretical B(EL), B(ML) values of inter-state transitions + energies of γ rays \rightarrow Distribution of γ -ray branches from state A to states B_i, already implemented in most calculations

Uncertainties also calculated via $\Delta |E_i - E_f|$ perturbation, and variations in effective charges

Combine parts A and B by funneling theoretical β and γ branches down to the ground state, with proper uncertainty propagation

γ-ray intensity comparison, ⁹⁹Cd



General agreement of the strongest γ -ray intensities from β decays from SM calculations:

- ⁸⁸Sr core (Z = 38, N = 50)
- Proton $p_{1/2}$, $g_{9/2}$ and neutron $d_{5/2}$, $g_{7/2}$, $d_{3/2}$, $s_{1/2}$, $h_{11/2}$ orbitals as valence space

No significant disagreements of the weaker γ -ray branches

βγ spectra (¹⁰¹Sn \rightarrow ¹⁰¹In), literature



K. Straub, PhD thesis (TU Munich, 2010) from GSI RISING campaign for ¹⁰⁰Sn, fragment separation for PID; much cleaner separation!



$\beta\gamma$ spectra (¹⁰¹Sn \rightarrow ¹⁰¹In), this work



Affirmed two γ rays at 1346 and 1500 keV and two new γ rays at higher energies Inconclusive on many other transitions, 252-keV γ belongs to the granddaughter ¹⁰¹Cd [M. Huyse et al., ZPA 330, 121 (1988)]

Statistically challenging even more than the prized ¹⁰⁰Sn, despite having been produced at least 4 times more... Why?

Combination of:

- Non-negligible β -delayed proton emission branch (~20%) to ¹⁰⁰Cd instead
- Fragmented β-decay branches
- Absence of low-energy γ rays, where EURICA efficiency is reasonably high

Reproducing βp spectrum from theory



For states with $E_x > S_p$, γ -decay competes with proton emission

Robust theory should be able to reproduce the experimental βp spectrum and branching ratio of ¹⁰¹Sn, which also affects I_v calculation

Theories on proton emission by Delion, Liotta and Wyss [Phys. Rep. 424, 113 (2006)] predict partial half-lives as a function of proton energy and angular momentum, etc.

In gds model space above N = Z = 50, dominant emission from g(l = 4) or d(l = 2) orbitals [P. J. Davies et al., PLB 767, 474 (2017)]

Branching ratios predicted within ~50%, distribution at higher energies missing 14

γ-ray intensity comparison, ¹⁰¹In



Strongest γ -ray intensities reproduced by theories, taking into account βp branching

SM-A: $J^{\pi}(^{101}Sn) = 5/2^+$, SM-B: $J^{\pi}(^{101}Sn) = 7/2^+$ assumptions to probe ground state dependence on γ -ray distribution

Some differences, but not convincing enough to accept or reject either scenario; slight favor to the $7/2^+$ case based on non-observation of low-energy γ 's

Proposed level scheme of ¹⁰¹In



 $4274 - - (7/2^+)$

2157

3259—(21/2⁺)

4074—7/2⁺ 2157-2116 cascade is tentative, only based on SM (coincidence predicted)
 3114—21/2⁺ Multiple spin candidates for 1346, 1500-keV gamma rays formed by 2-neutron configurations



Still very uncertain

above N = 50 shell

Summary

Extending theory to compare to limited experimental results in exotic decay spectroscopy of ⁹⁹In and ¹⁰¹Sn

- From B_{GT} distribution to βp energy spectra
- From excitation energies to γ-cascade schemes
 → infer J^π from theory and γγ coincidence relationships
- Theoretical uncertainty estimation crucial

Key requirements on studying the rarest isotopes, when statistics are limited:

- High-purity data \rightarrow high suppression of background and contaminants
- High-efficiency detector systems $\rightarrow \gamma\gamma$ coincidences, for instance
- Capability to distinguish multiple excitation and decay/exit channels: β vs βp, βn, β2n, etc